

The power of teaming-up HST and Gaia

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Dynamical tracers to investigate dark matter haloes



Dynamical tracers
(radial velocities)

- Mass
- Shape

Dynamical tracers to investigate dark matter haloes



Dynamical tracers
(radial velocities)

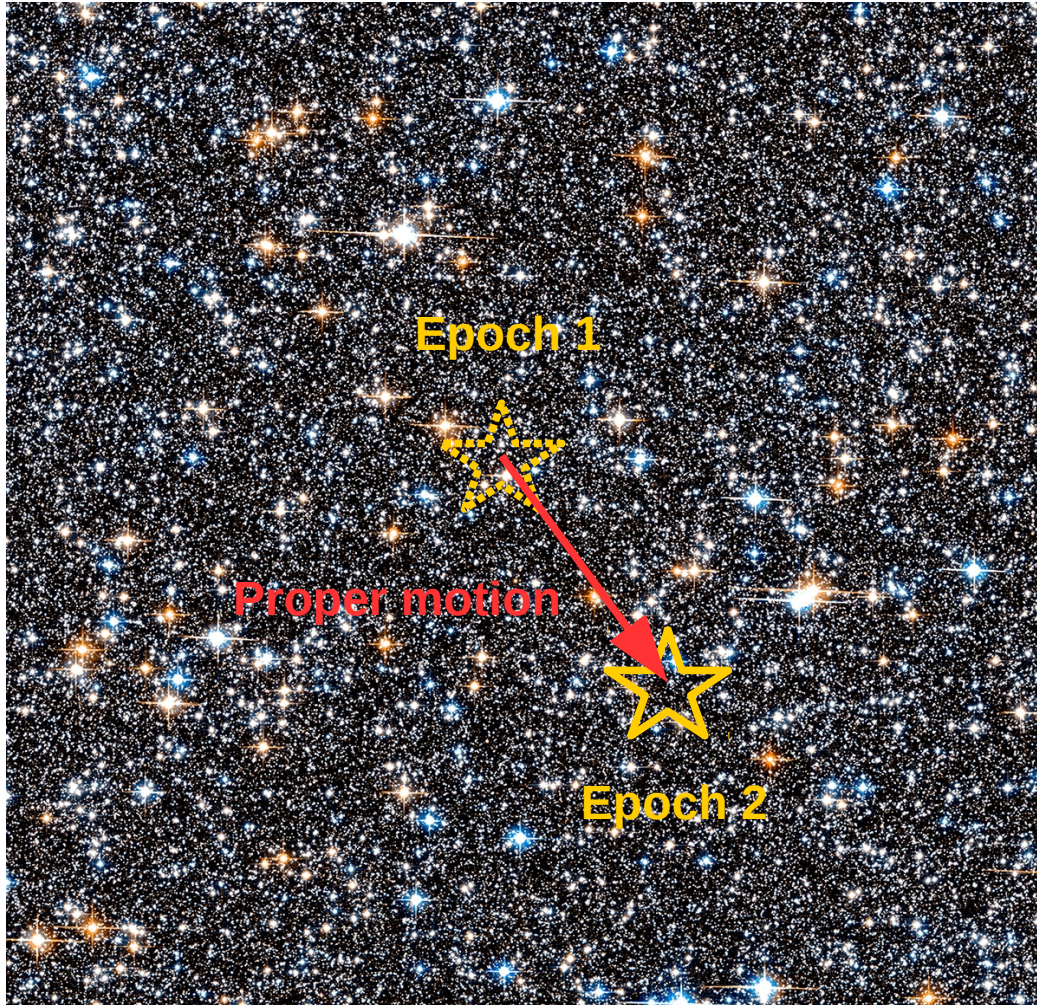
- Mass
- Shape

MISSING



PROPER MOTIONS

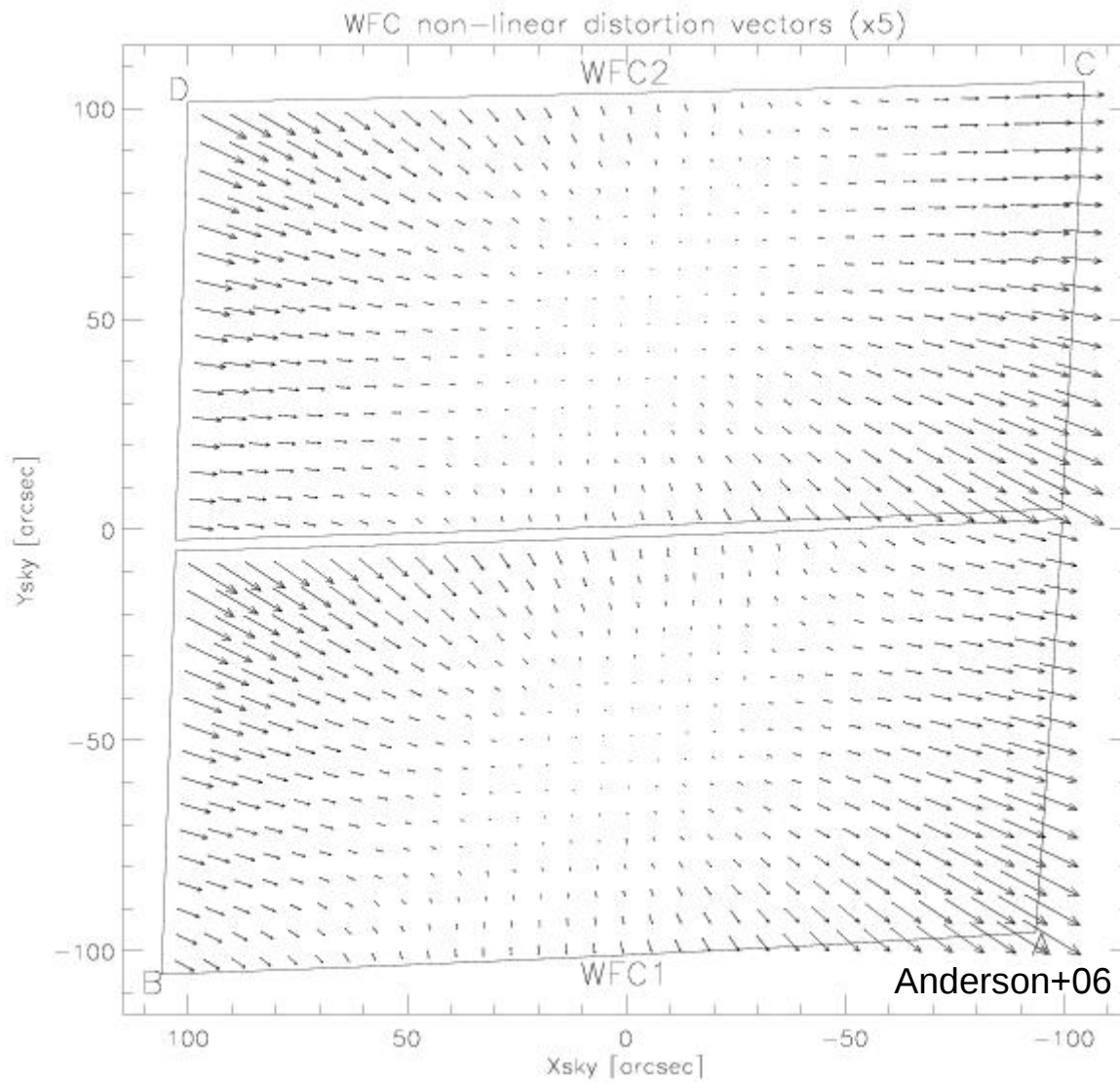
Proper Motions



Very complex measurement!

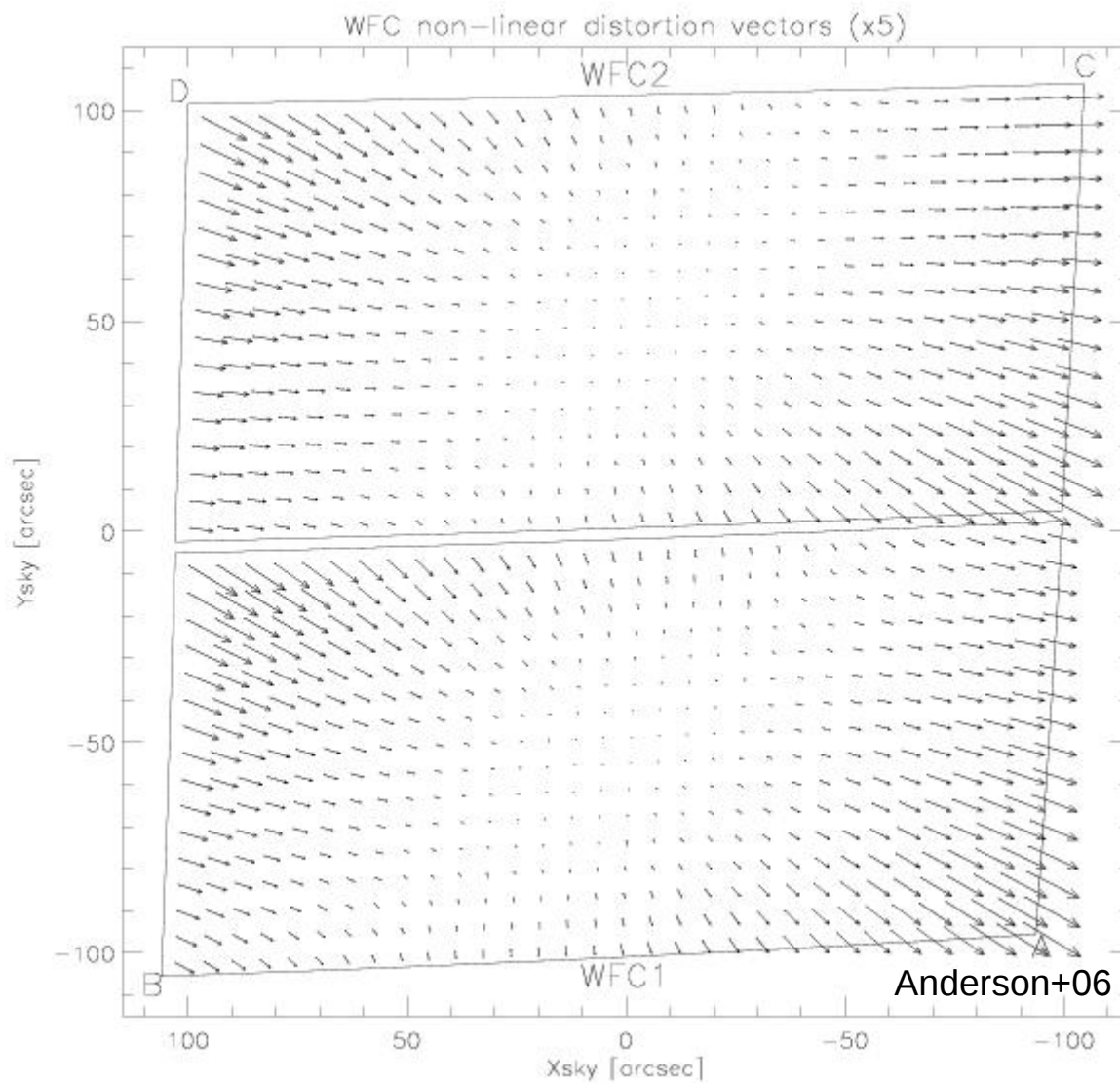
- 1- At least two observations
- 2- Short separation=small the proper motion
- 3- Large distance=small the proper motion
- 4- Instrumental effects >>> proper motions

Proper Motions: HST



PSF knowledge 0.01 pixels

Proper Motions: HST



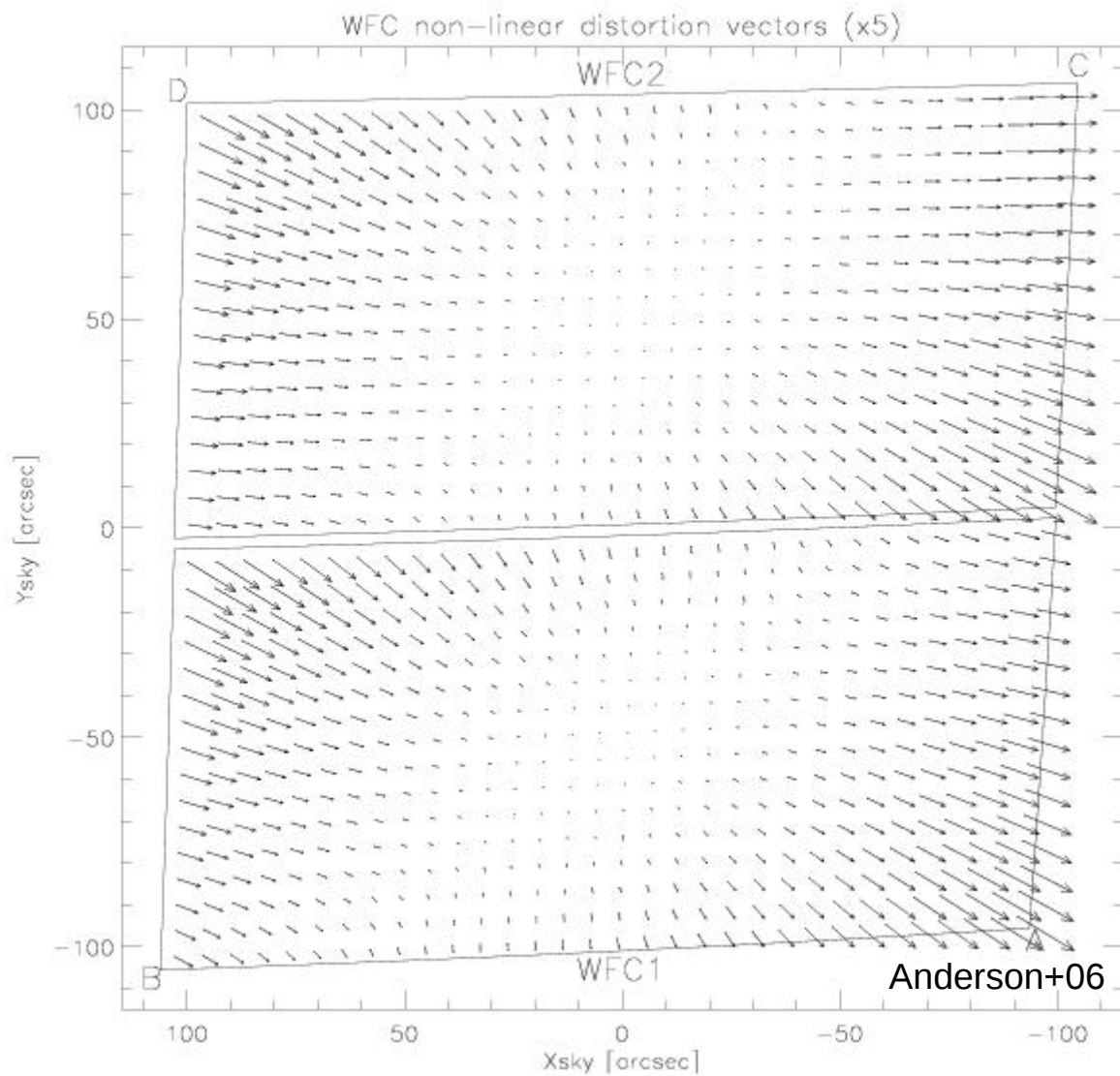
PSF knowledge 0.01 pixels

+

+

Geometric distortion corrections 0.01 pixels

Proper Motions: HST

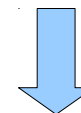


PSF knowledge 0.01 pixels

+

+

Geometric distortion corrections 0.01 pixels

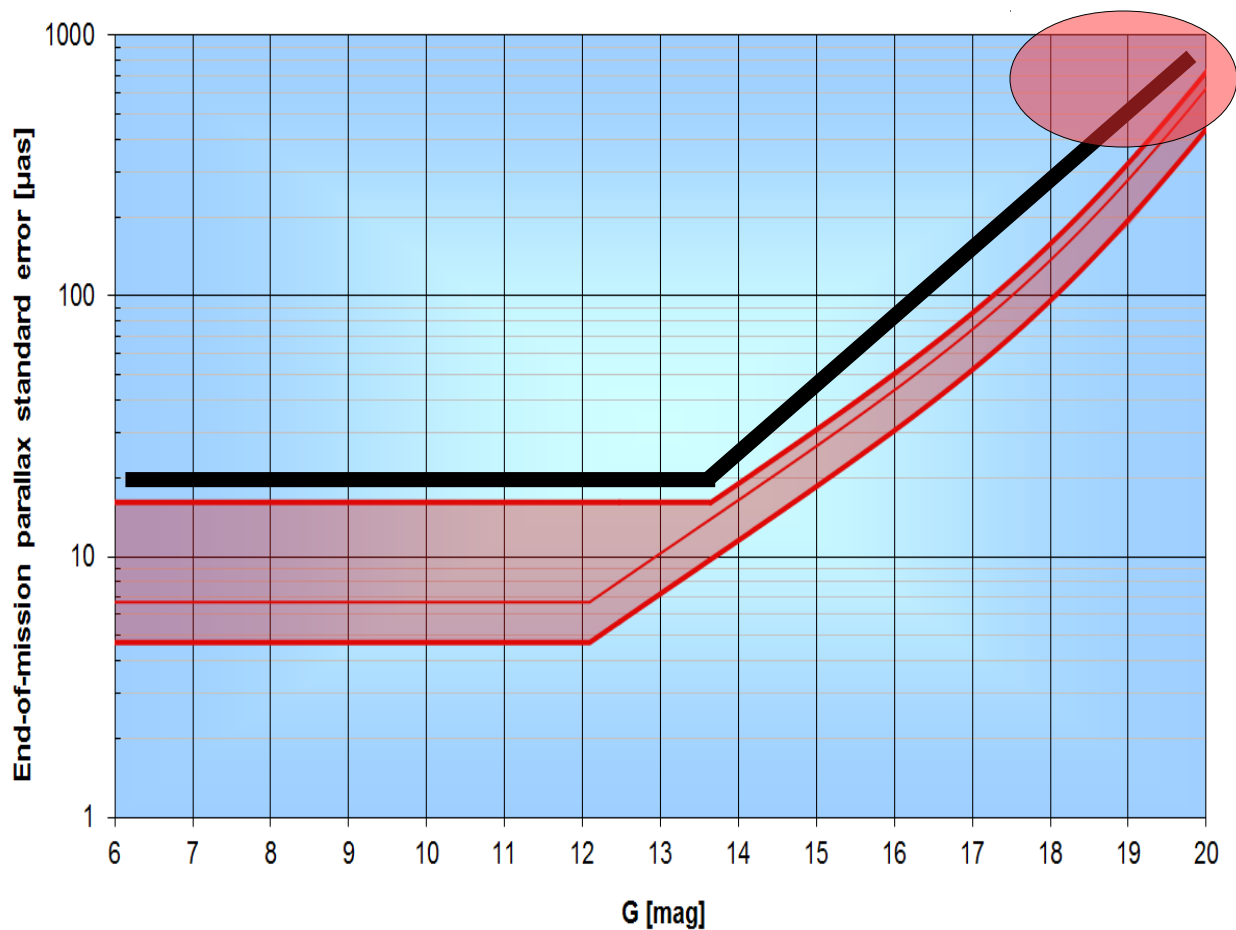


Absolute positional precision

=

0.7 mas

Proper Motions: *Gaia*

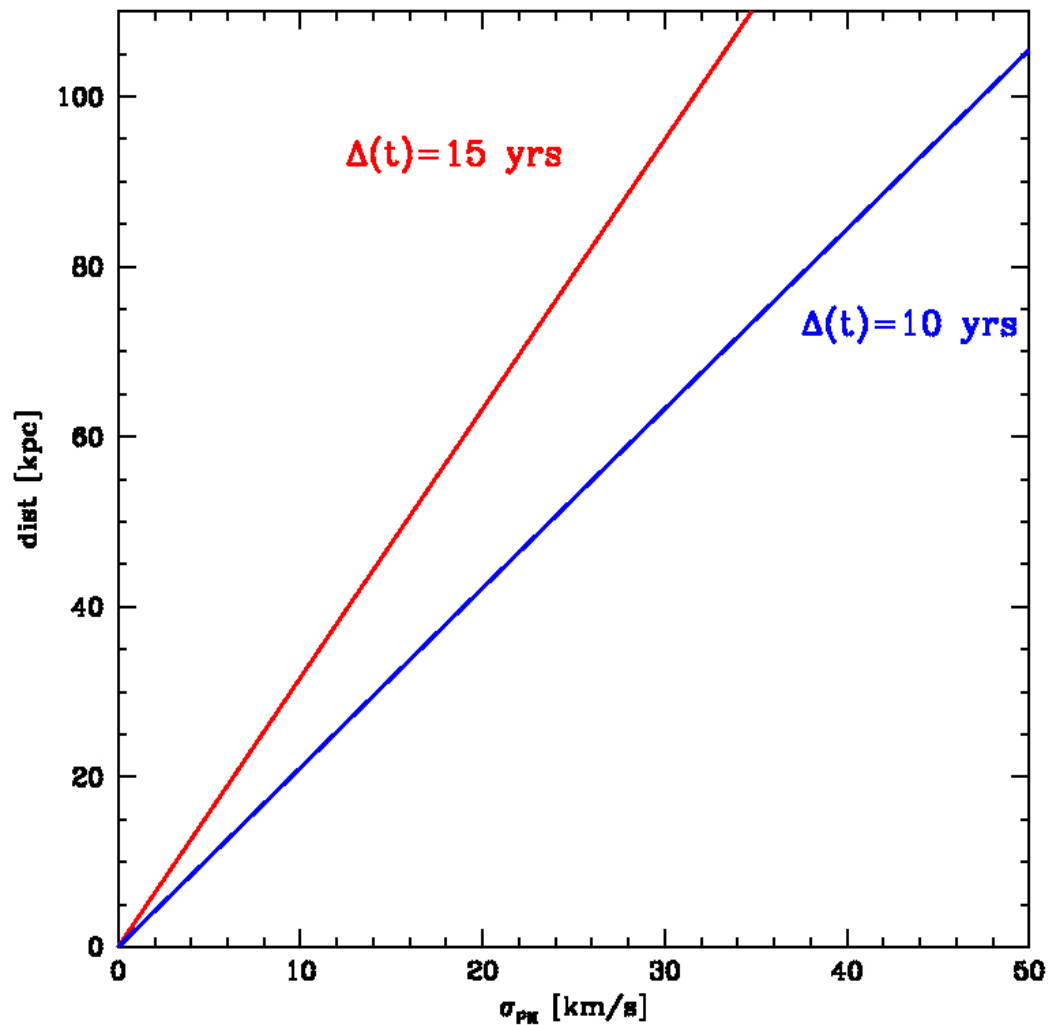


At the typical magnitude of
HB for 100 kpc distant
stellar population

=

Precision similar to HST!

Proper Motions: HST + *Gaia*



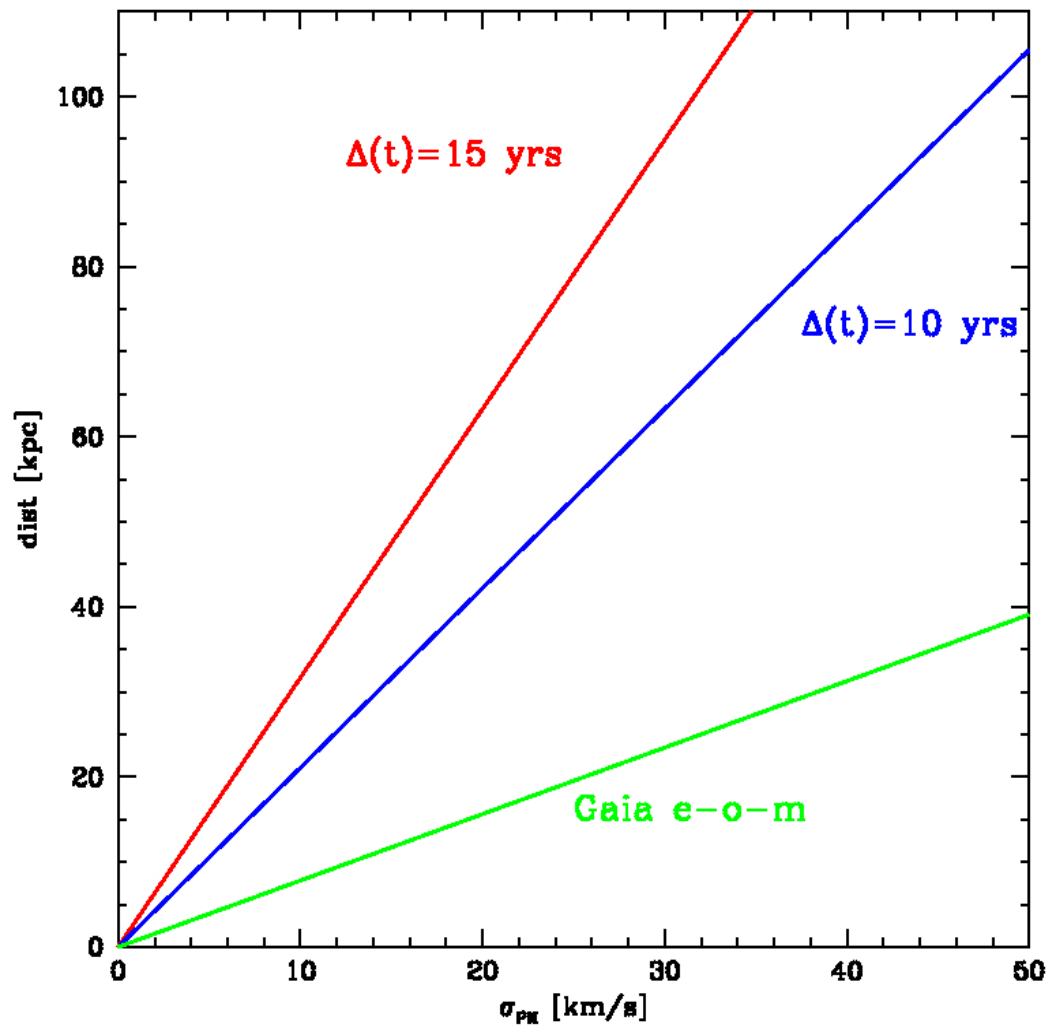
Absolute proper motion precision

< 50 km/s

Gaia end of mission: 120 km/s!

Gain coming from Δt

Proper Motions: HST + *Gaia*



Absolute proper motion precision

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Proof of concept: NGC 2419

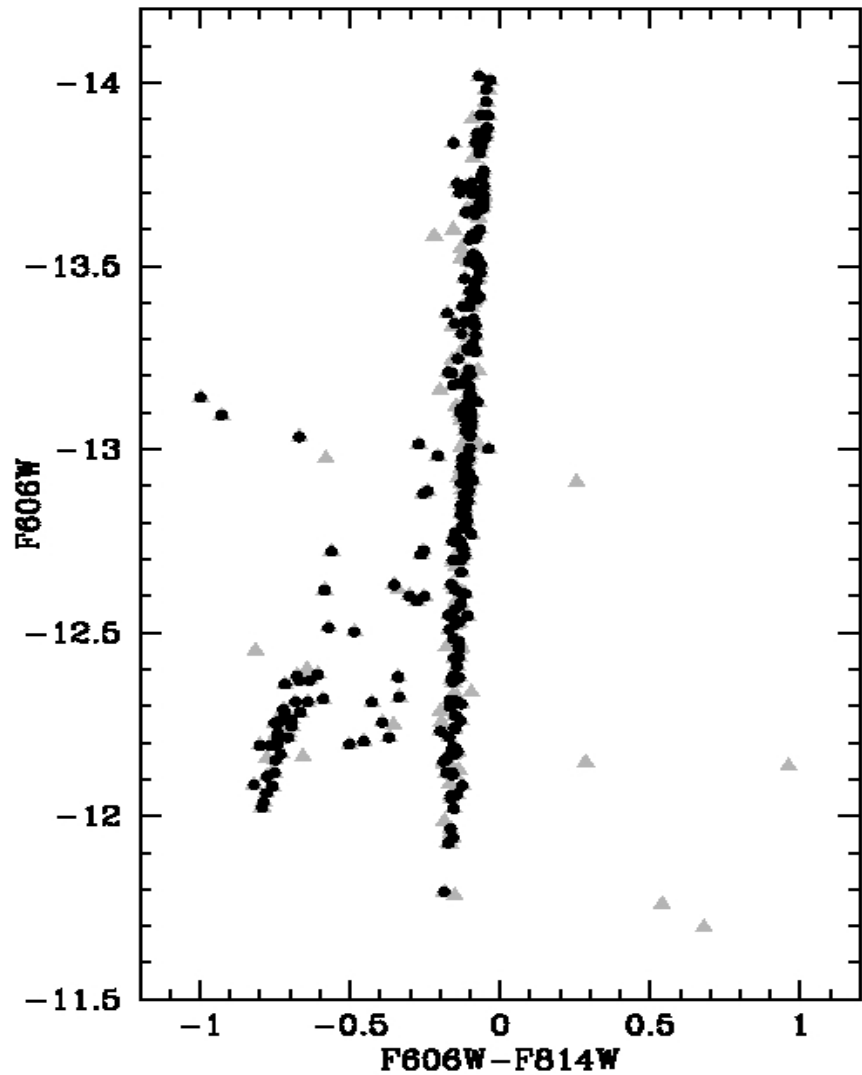
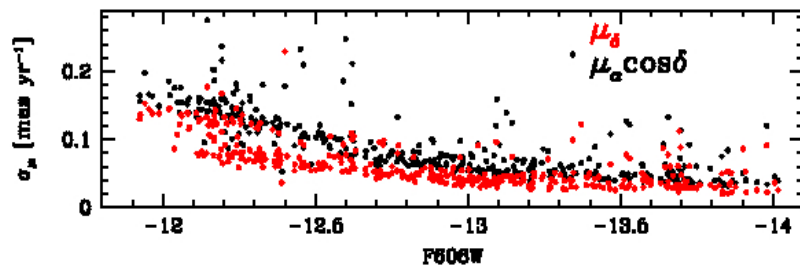
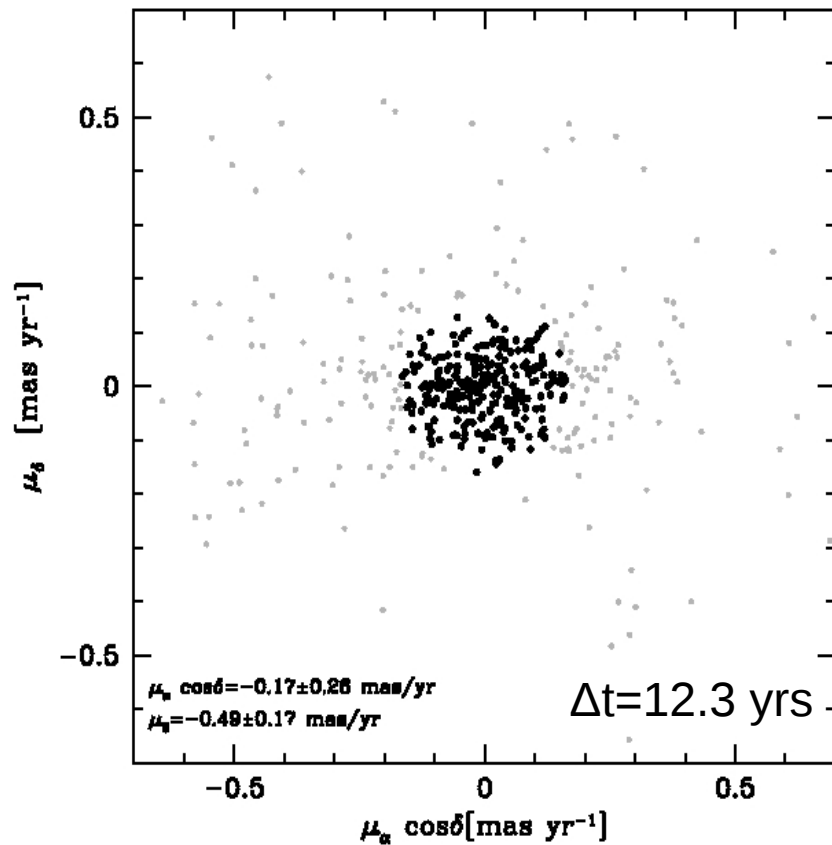


- Distance = 87.5 kpc
(Di Criscienzo+11)
- Brightest halo GC
(Harris+96)
- Very extended
(Harris+96)
- Ca, K spread
(Mucciarelli+12, Lee+13)

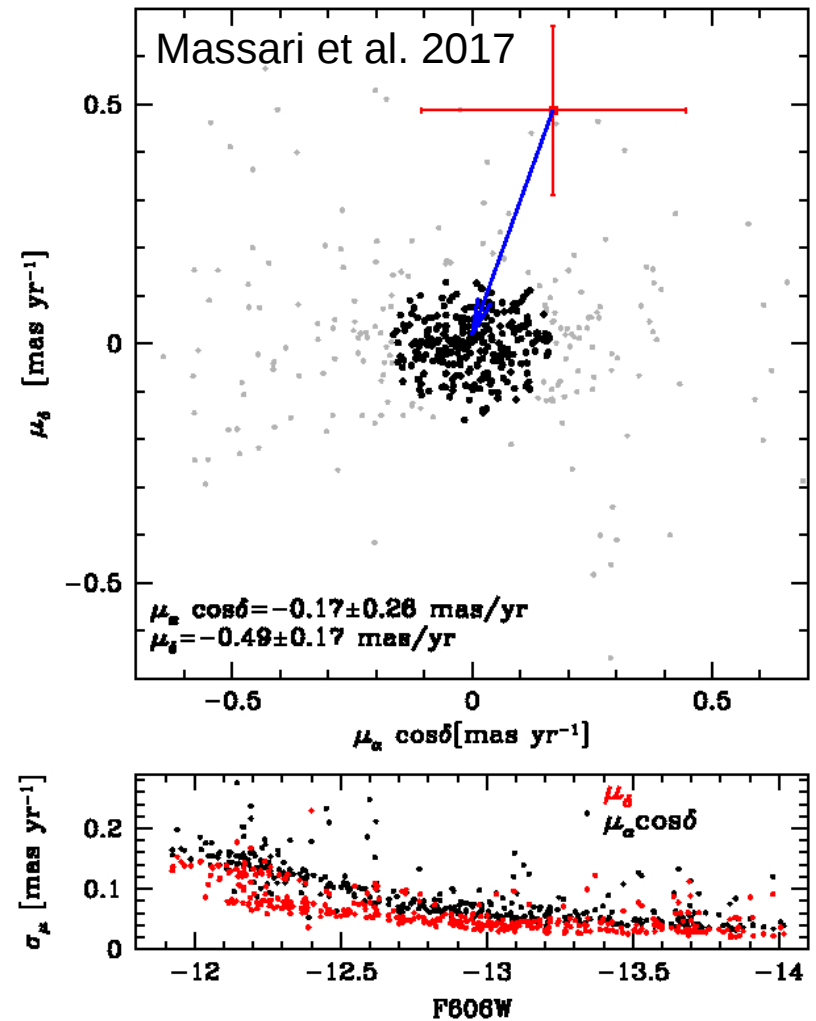
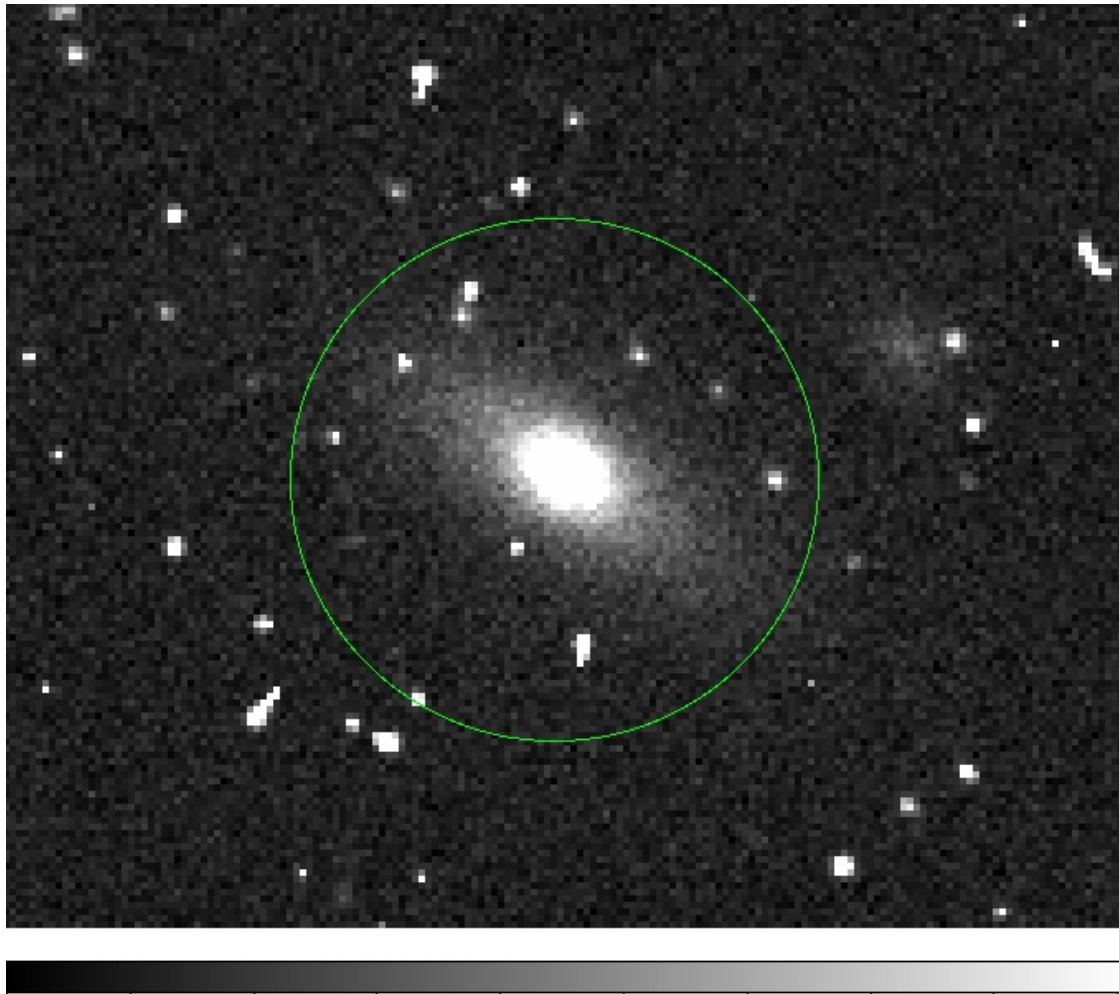
EXTERNAL ORIGIN?
(Irwin+99, Belokurov+14)

Proof of concept: NGC 2419

CataPack code for geometric transformations

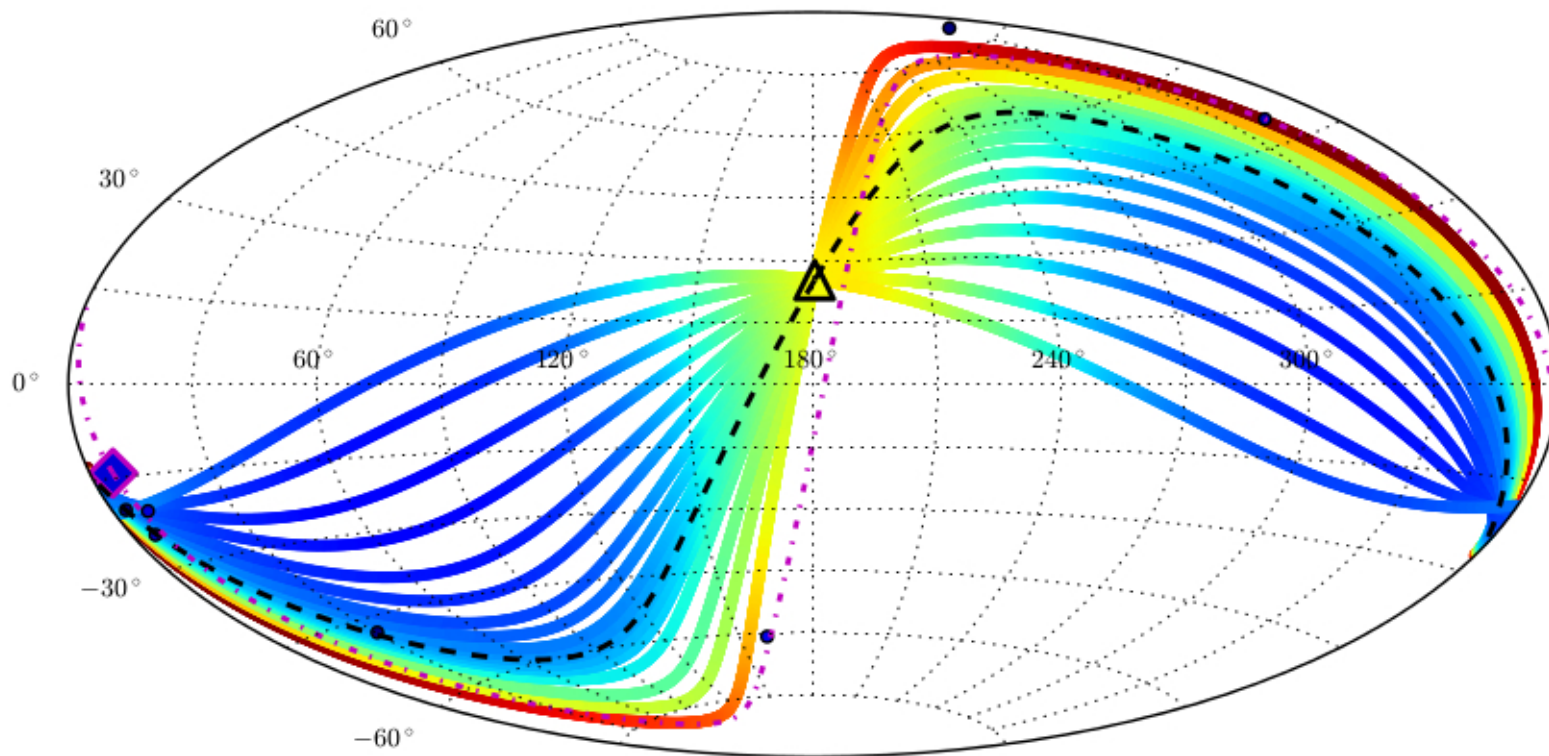


Proof of concept: NGC 2419

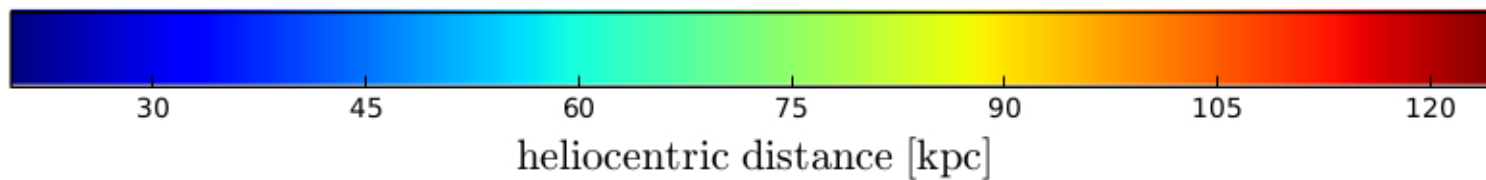


$$(\mu_\alpha \cos(\delta), \mu_\delta) = (-0.17 \pm 0.26, -0.49 \pm 0.17) \text{ mas/yr}$$

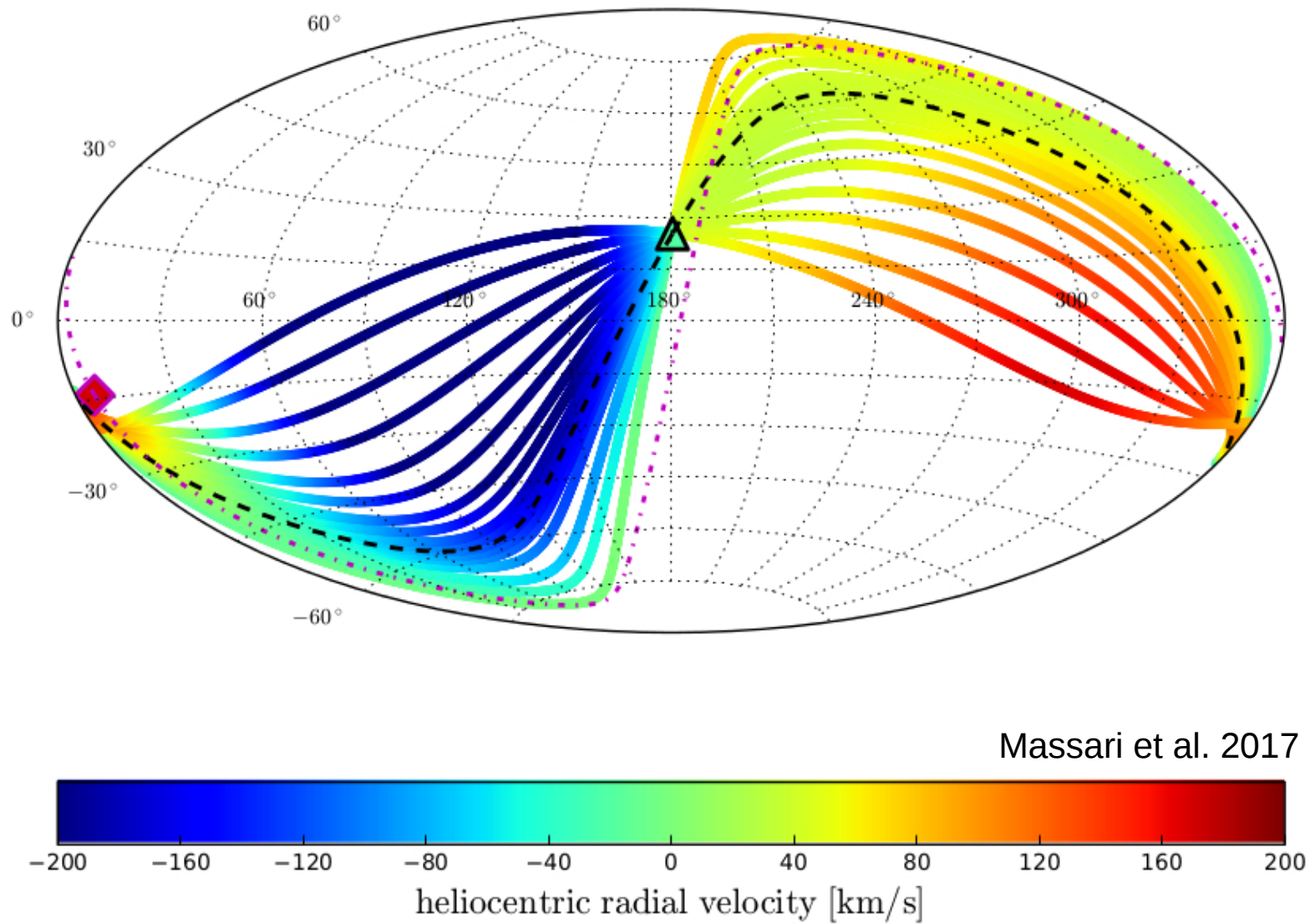
Proof of concept: NGC 2419



Massari et al. 2017



Proof of concept: NGC 2419



NGC 2419 and Sgr dSph

Features of the orbit:

- close to polar



NGC 2419 and Sgr dSph

Features of the orbit:

- close to polar



- clockwise



NGC 2419 and Sgr dSph

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- $r_{\text{apo}}=98\pm 2$ kpc

- $r_{\text{peri}}=53\pm 24$ kpc

NGC 2419 and Sgr dSph

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If the progenitor had larger mass, then debris predicted at same distance (Helmi & White 2001)

NGC 2419 and Sgr dSph

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- BHB tidal stream overlapping in position and l.o.s velocity: part of Sgr trailing stream (Belokurov+2014)

- Model by Vera-Ciro+13, Sgr trailing stream with tangential velocity similar to NGC2419!

NGC 2419 and Sgr dSph

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Likely association!

Summary

- HST + Gaia: high accuracy down to mag 20 → HB of stellar populations 100 kpc distant
- Test on the distant GC NGC2419 (no PM so far)
- Presence of 1 background galaxy → $(\mu_{\alpha\cos(\delta)}, \mu_{\delta}) = (-0.17 \pm 0.26, -0.49 \pm 0.17)$ mas/yr
- Computation of the orbit: Likely association between NGC2419 and the Sgr dSph

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Next steps

- Try this strategy on dwarf galaxies
- First target: Sculptor dwarf spheroidal (debated PM measurement)